



From the Big Bang to the Nobel Prize and the End of the Universe

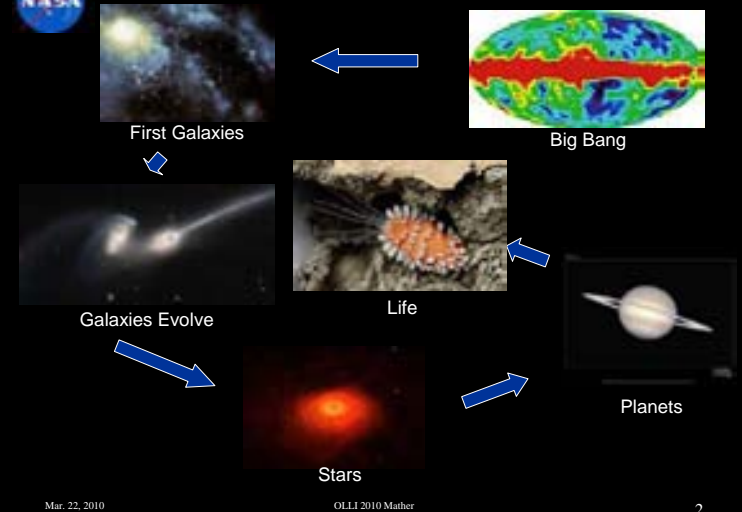
John C. Mather
Senior Project Scientist, James Webb Space Telescope,
NASA's Goddard Space Flight Center
Mar. 22, 2010

Mar. 22, 2010

OLLI 2010 Mather

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Astronomical Search For Origins



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Can you imagine?
**Your chin is made
of exploded stars!**

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Looking Back in Time

HAND		1 m	0.000 000 003
EARTH		7000 km	0.0 2 SEC
SUN		150,000,000 km	500 CE
STAR			4 YRS
GALAXY			25,000 YRS
BIG BANG	?		15,000,000,000 Y

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Measuring Distance

Earth NOW
1 AU
Sun
6 Months Later
 p = parallax (angle)
 d = distance

1. TRIANGLES

KNOWN LENGTH
KNOWN ANGLE
STAR

This technique enables measurement of enormous distances

2. STANDARD CANDLES

$\frac{\text{BRIGHTNESS}_1}{\text{BRIGHTNESS}_2} = \frac{r_2^2}{r_1^2}$

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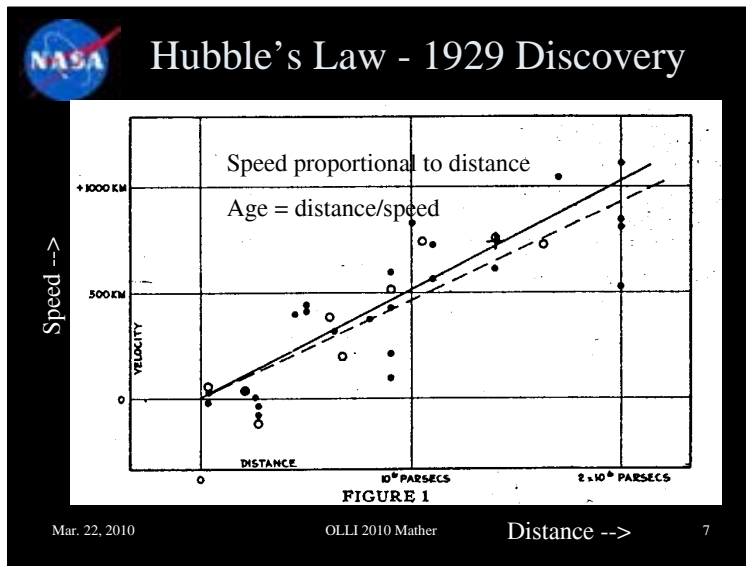
Astronomer's Toolbox #2: Doppler Shift - Light

Continuous Spectrum
Emission Line Spectrum
Absorption Line Spectrum

OBJECT RECEDING: LONG RED WAVES
OBJECT APPROACHING: SHORT BLUE WAVES

400 500 600 700

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The Power of Thought

Alexander Friedman

Georges Lemaitre & Albert Einstein

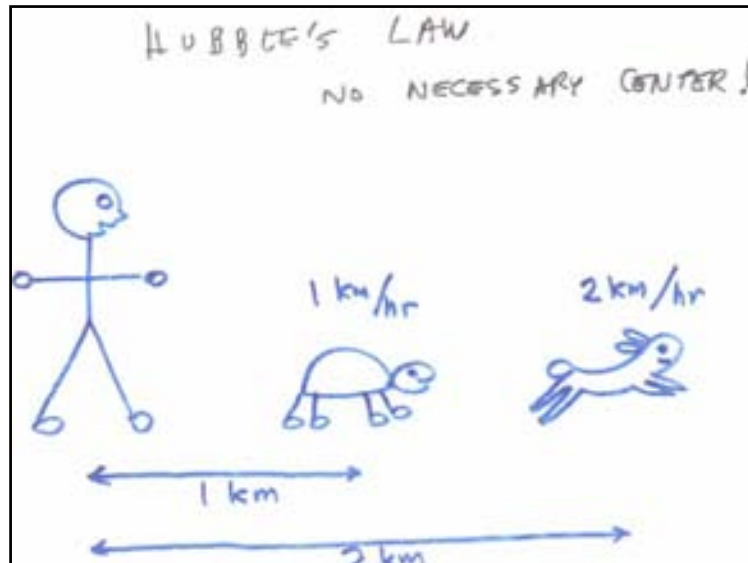
George Gamow

Robert Herman & Ralph Alpher

Rashid Sunyaev

Jim Peebles


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Big Bang - Cosmic Explosion 13.7 billion years ago

IMPOSSIBLE TO DRAW A PICTURE!


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So what happened?

- Primordial material, possibly infinite in every dimension
- Small piece of it (10 cm in size?) does something quantum mechanical with unknown physics
- Rapid expansion, faster than light can keep up with, stretches this little bit into whole observable universe (cosmic inflation)

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How did the whole observable universe fit into that little ball?

- Space is mostly empty - stars are very very far apart
- Atoms are mostly empty - atomic nuclei are very tiny compared to size of atoms
- Squeeze very hard, and compression can create antimatter and rip quarks apart inside protons and neutrons
- Higher temperatures fit more particles into given volume
- Inflation turns “false vacuum” energy into real particles
- Squeeze even harder, and the known laws of physics no longer apply - space and time may mix into higher dimensions?

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How did a smooth Big Bang make complicated things like us?

- Gravity is long range attractive force
 - Matter distribution is unstable
 - Remove heat, and system heats up more
 - Makes condensed objects (stars, galaxies, etc.)
 - Gravitational energy flows support complexity
- Stars release heat from nuclear reactions
 - Heat & light received by Earth support complexity, from weather to photosynthesis

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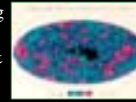
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Early History of the Universe

Big Bang
seen by
COBE &
WMAP



?

Galaxy
assembly



?

Galaxies,
stars,
planets,
life.



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- Horrendous Space Kablooney - exponential expansion, primordial fluctuations, matter/antimatter, dark matter, dark energy, 13.75 ± 0.11 billion years ago (WMAP7)
- Annihilation of antiparticles, 1 part per billion matter remaining
- Formation of Helium nuclei, 3 minutes, redshift $z = 10^9$
 - $[1+z = \text{size of universe now} / \text{size then}]$
- Formation of neutral gas “recombination”, 389,000 yrs, $z=1089$
- Population III supermassive stars, super-supernovae, and black holes, $z=17$ (age ~ 200 Myr)
- Galaxy formation in small parts, star formation, merging and clustering of galaxy parts, until $z \sim 1$
- Expanding universe begins to accelerate, 5 billion years ago

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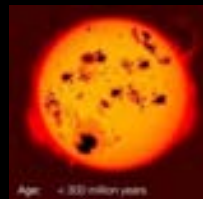
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Possible Early History of Earth

- Sun and first solid bodies in Solar System 4.567 billion years ago
- Mars-sized body “Theia” hits Earth, melting everything, dispersing volatiles like C and H; debris forms Moon, 90 MY AF (after formation)
- Cool early Earth, possibly with water
- Jupiter, Saturn orbits switch twice, clear debris from solar system, cause “late heavy bombardment”, “Hadean” geologic period, many craters, new water and carbon delivery to Earth, 400 - 700 MY AF
- Life forms shortly after (~ 3.8 BY ago); all life uses same genetic code! (we are related to bugs)
- Young Sun very active, gets steadily brighter with time, warming Earth

Hartman & Davis idea 1975



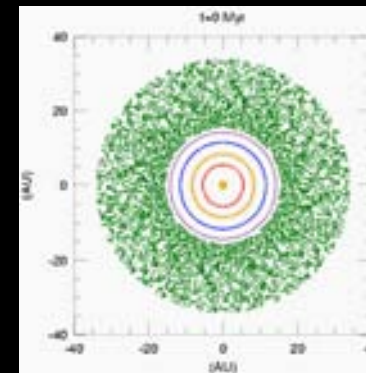
Age: ~ 300 million years
Ed Guinan 2009

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Possible Early Solar System



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Continents Floating and Moving

- Huge volcanic effects on atmospheric composition (CO₂, H₂S, etc. fluctuate)
- Vaalbara, 3.3-3.6 billion years ago
- Rodinia, 1100 – 750 million years ago, split into proto-Laurasia, proto-Gondwana, and Congo Craton “Re-unite Gondwanaland”
- Pangaea, 250 MY ago
- Atlantic Ocean opens, ~ 100 MY ago



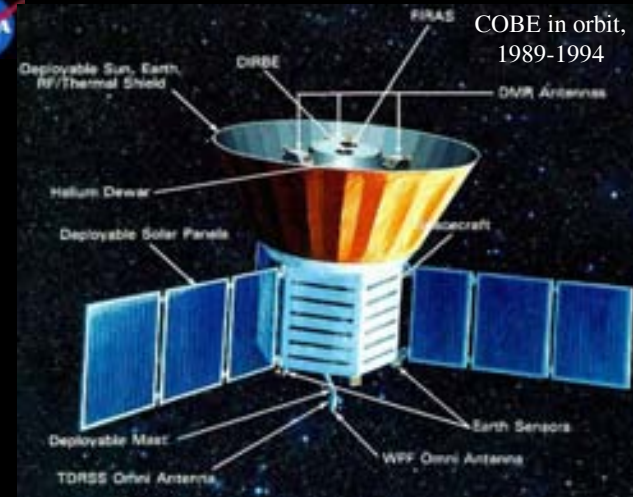
Ice Ages, Heat Ages, and Extinctions

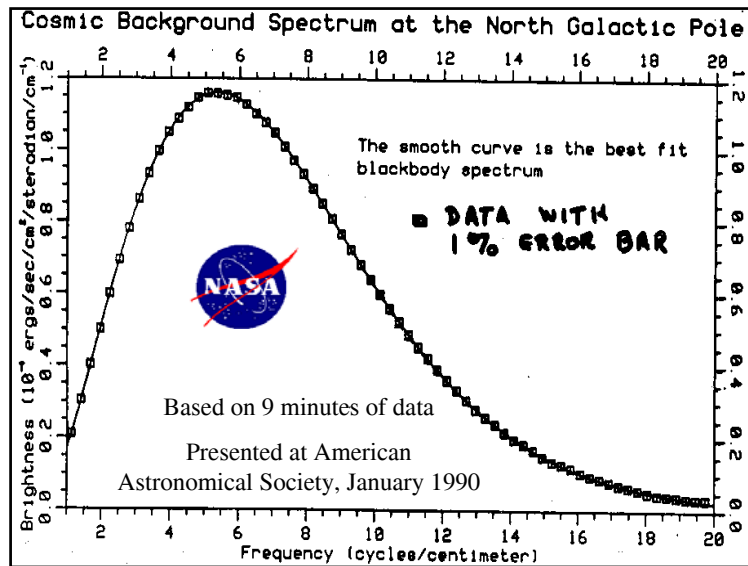
- Huronian Ice Age, 2.7 – 2.3 BYA
- Cryogenian, 850 – 630 MYA, possible “snowball Earth”
- Volcanism releases CO₂, warms Earth, enables Cambrian explosion of life, 530 MYA
- Andean-Saharan Ice Age, 460 – 430 MYA (an extinction event)
- Coal formation, 354 to 290 MYA
- Heat extinction event, 250 MYA, life retreats to Antarctica
- Dinosaur extinction, coincides with (multiple?) meteor events, 65 MYA
 - Chicxulub impact on the Yucatan Peninsula off Mexico?
 - Volcanic activity in Deccan Traps in India?
- Current Ice Age, 2.58 MYA, ending 10,000 YA
 - Riss, 180,000 – 130,000 YA (when Homo Sapiens developed in Africa?)
 - Wurm, 70,000 – 10,000 YA (begin modern civilization)



Civilization and the Future

- Galileo’s telescope 1609 (2009 International Year of Astronomy)
- 2012, nothing happens!
- Possible far future: all the CO₂ goes into rocks, BIG FREEZE
- 1 BY in future, sun is brighter, Earth gets too hot for us
- 5 BY, sun becomes red giant, and Andromeda Nebula collides with Milky Way
- 7.6 BY, sun goes out
- Accelerating universe continues, galaxies recede, stars go out: dark!





PHYSICS TODAY

JUNE 1992

Sky map from DMR, 2.7 K +/- 0.003 K

Doppler Effect of Sun's motion removed ($v/c = 0.001$)

Cosmic temperature/density variations at 389,000 years, +/- 0.00003 K (part in 100,000)

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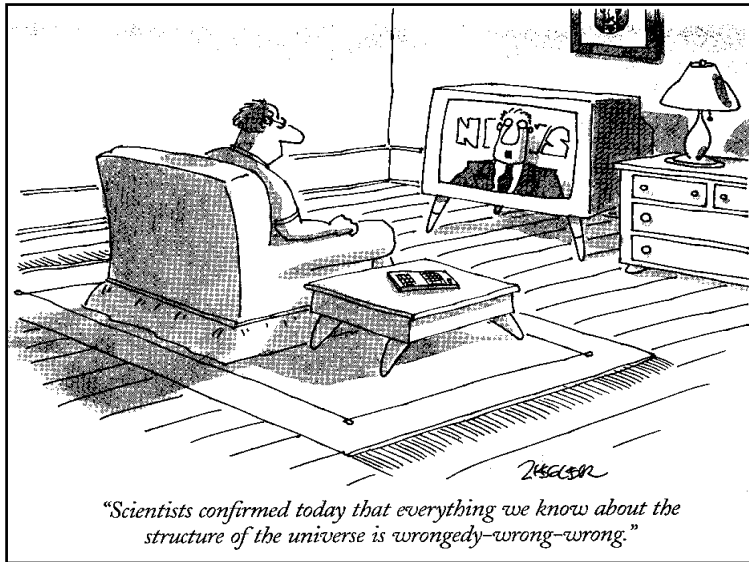
Nobel Prize Press Release


The Royal Swedish Academy of Sciences has decided to award the Nobel Prize in Physics for 2006 jointly to **John C. Mather**, NASA Goddard Space Flight Center, Greenbelt, MD, USA, and **George F. Smoot**, University of California, Berkeley, CA, USA "for their discovery of the blackbody form and anisotropy of the cosmic microwave background radiation".

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From Press Conference to Stockholm



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Dark Energy!


MacArthur Fellow
2008 - Adam Riess

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S. Perlmutter, A. Riess, B. Schmidt




A few big mysteries...

- Why is there matter and no antimatter?
- What is dark matter?
- What is dark energy?
- Was Einstein right about relativity?
- How did we get here?
 - Formation of stars, chemical elements, galaxies, planets, ...
- Are we alone?
 - How did Earth become habitable?
 - Any other places that could support life?
- What happens next?

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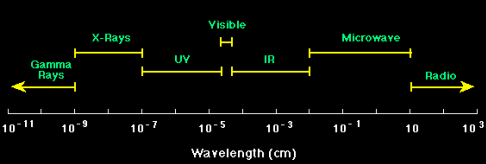
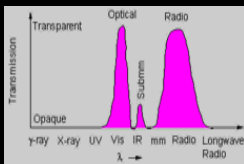
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Light comes in more colors than our eyes can see

Light from the first galaxies is **redshifted** from the visible into the infrared.

Infrared is heat radiation

Our eyes can't see it, but our skin can feel it

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NASA James Webb Space Telescope (JWST)

Organization

- Mission Lead: Goddard Space Flight Center
- International collaboration with ESA & CSA
- Prime Contractor: Northrop Grumman Aerospace Systems

Instruments:

- Near Infrared Camera (NIRCam) – Univ. of Arizona
- Near Infrared Spectrograph (NIRSpec) – ESA
- Mid-Infrared Instrument (MIRI) – JPL/ESA
- Fine Guidance Sensor (FGS) – CSA

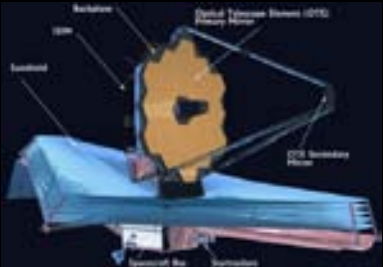

Operations: Space Telescope Science Institute

Description

- Deployable infrared telescope with 6.5 meter diameter segmented adjustable primary mirror
- Cryogenic temperature telescope and instruments for infrared performance
- Launch June 2014 on an ESA-supplied Ariane 5 rocket to Sun-Earth L2
- 5-year science mission (10-year goal)

www.JWST.nasa.gov

JWST Science Themes

End of the dark ages: First light and reionization

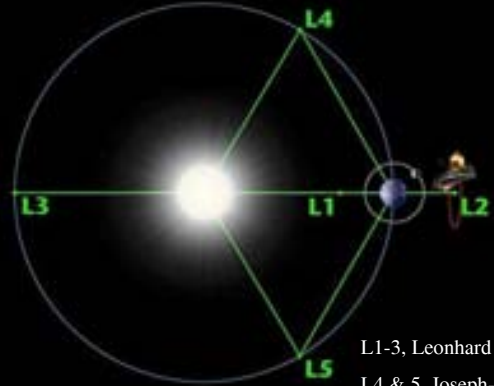
The assembly of galaxies

Birth of stars and proto-planetary systems

Planetary systems and the origin of life

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NASA JWST Orbits the Sun-Earth Lagrange Point L2



L1-3, Leonhard Euler, 1750.
L4 & 5, Joseph-Louis Lagrange, 1772

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NASA Northrop Grumman's JWST model

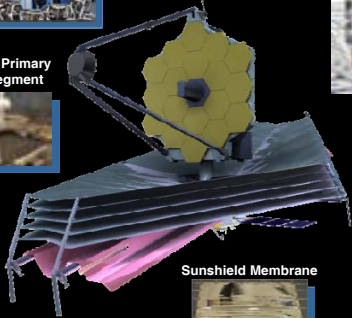
Washington, DC 2007

Munich, Germany 2008



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NASA JWST Technology



Backplane

Mirror Phasing Algorithms

Beryllium Primary Mirror Segment

Sunshield Membrane

Near-Infrared Detector

Mid-Infrared Detector

Cryogenic ASICs

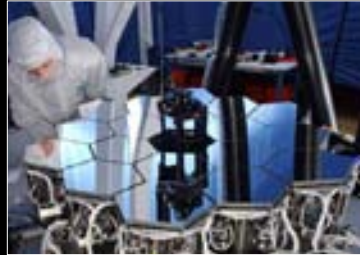
μShutters

Cryocooler

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Testbed Telescope



- 1/6 scale model with all the same adjustments
- Proves that all the adjustment procedures work as expected

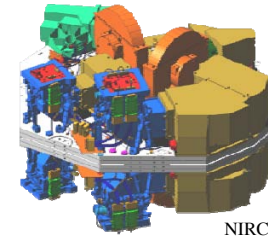
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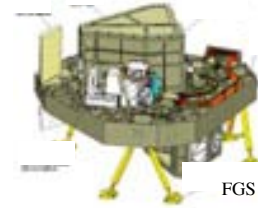
Four science instruments enable imagery and spectroscopy over the 0.6 – 29 micron spectrum



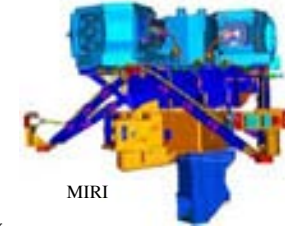
NIRCam



NIRSpect



FGS



MIRI

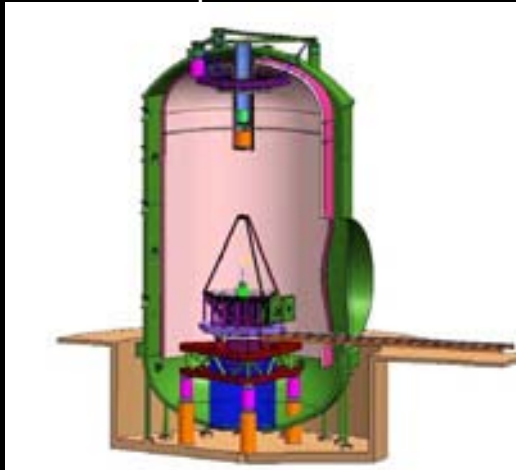
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JWST cold optical test in Houston



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Nature's Lenses in Space



Galaxy Cluster Abell 370
Hubble Space Telescope + ACS/WFC

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NASA, ESA, the Hubble 30th BPO Team, and ST-ECF

STScI-PR09-239

NASA Dwarf Galaxies with Dark Matter

Perseus Cluster Dwarf Galaxies
Hubble Space Telescope • ACS/WFC

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NASA, ESA, and C. Conselice, University of Nottingham. ©ESA/ESA/HST

NASA How do galaxies evolve?

Interacting Galaxies Hubble Space Telescope • ACS/WFC • WFC3

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NASA, ESA, the Hubble Heritage (STScI/AURA)-ESA/Hubble Collaboration and the Johns Hopkins University, University of Wisconsin-Madison, Ohio State University. ©ESA/ESA/HST

NASA Where and when did the Hubble Sequence form?
How did the heavy elements form?

- Galaxy assembly is a process of hierarchical merging
- Components of galaxies have variety of ages & compositions
- Observations:
 - NIRCam imaging
 - Spectra of 1000s of galaxies

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NASA SN 2006 gy – brightest supernova

- Could be the first observation of a pair-production instability, from the death of a very massive star.
 - Stars are normally held up by the balance of light pressure and gravity
 - Gamma rays producing electron/positron pairs scatters light, reducing pressure. Instability creates runaway collapse.
- A nearby analog for the first stars in the Universe.

$\gamma \rightarrow e^- + e^+ \rightarrow \gamma$

- Progenitor was similar to Eta Carina. Hubble Image of Eta Carina.

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NASA

Gamma Ray Burst 4/23/09 was one of the most distant objects yet found ($z = 8.2$) – supernova jet aimed at us!

Number of Gamma-Ray Bursts

Age of the Universe: billions of years

Redshift

GRB 090423 (z=8.2)

JANUS GRB (SMEX) search proposed, could see to $z = 12$

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NASA

How stars and planets form?

Deeply embedded protostar Circumstellar disk

10^4 yrs; $10-10^4$ AU; $10-300$ K

10^5 yrs; $1-1000$ AU; $100-3000$ K

10^{6-7} yrs; $1-100$ AU; $100-3000$ K

10^{7-8} yrs; $1-100$ AU; $200-3000$ K

Agglomeration & planetesimals Mature planetary system

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NASA

The Eagle Nebula as seen with Hubble

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NASA

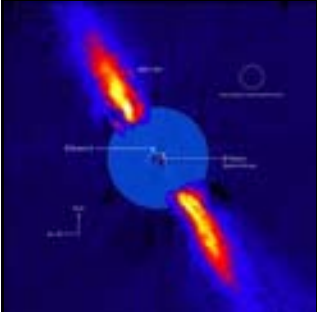

The Eagle Nebula as seen in the infrared

M. J. McCaughrean and M. Andersen, 1994

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Planets Seen!

HR 8799 b,c,d
Marois et al. 2008
Gemini & Keck

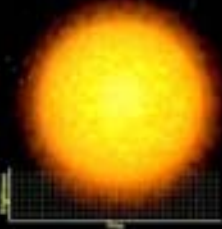




β Pictoris b
A.-M. Lagrange et al. 2008
VLT

Fomalhaut b
Kalas et al. 2008
HST

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Primary **Secondary**

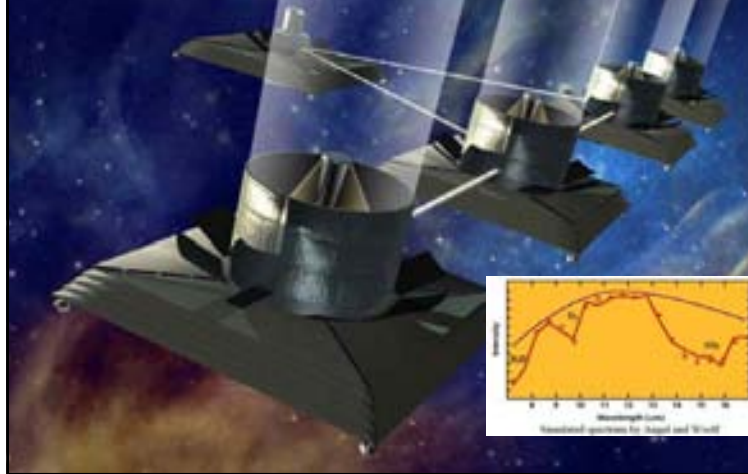
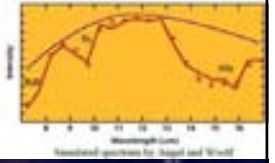
- Planet blocks light from star
- Visible/NIR light (Hubble/JWST)
- Radius of planet/star
- Absorption spectroscopy of planet's atmosphere
- JWST: Look for moons, constituents of atmosphere, Earth-like planets with water

- Star blocks light from planet
- Mid-Infrared light (Spitzer/JWST)
- Direct detection of photons from planet
- Temperature of planet
- Emission from surface
- JWST: Atmospheric characteristics, constituents of atmosphere, map planets

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Terrestrial Planet Finder Concept - Interferometer

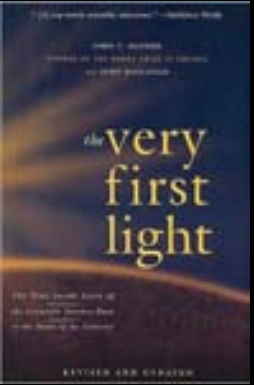



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 More Info:

- <http://www.jwst.nasa.gov>
- <http://lambda.gsfc.nasa.gov/>
- <http://nobelprize.org>
- Book, 2nd Edition:



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The End

And the beginning!

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 Rutgers University Lusscroft Farm - Site of Early Nerds in Sussex County, NJ



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JWST Deployment



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